

QUANTUM HARD-SPHERE ASSEMBLY OF FERMIONS

Chelimo S.L.¹, Achieng A.J.², Ayodo K.², Khanna K.M.¹, Tonui J.K.¹, Korir P.K.¹ and Kibet J.K.³

¹Department of Physics, University of Eldoret, P.O. Box 1125-30100, Eldoret-Kenya

²Department of Physics, Masinde Muliro University of Science and Technology, P.O. Box 190-50100, Kakamega-Kenya

³Department of Chemistry, Egerton University, P.O. Box 536-20115, Egerton- Kenya

*Author for Correspondence

ABSTRACT

The thermodynamic properties of a hard-sphere assembly of fermions in which the ^3He atoms are assumed to be interacting via a hard-sphere interaction has been calculated. The variation of energy per particle $\left(\frac{E}{N}\right)$ with temperature T is non-linear due to the thermal factor, but becomes insignificantly minimal at low temperatures below 4 K due to zero-point effects becoming visible. The specific heat C_v increases with increase in temperature to a point of inflection at a transition temperature $T_c = 18.67$ K and this value closely compares well with $T_c = 20.3$ K calculated using the Fermi energy formulation. This result indicates that, a hard-sphere Fermi gas of ^3He will undergo a phase transition at $T_c = 18.67$ K.

Keywords: Thermal Factor, Fermions, Transition Temperature

INTRODUCTION

Fermions have half-integral spin and constitute the second half of the particles family. To tell whether an atom is a fermion we look at the total number of protons, neutrons, and electrons making up that atom. Since these are all spin $\frac{1}{2}$ fermions adding up an odd number of them will make an atom a fermion (half integer spin). Fermions obey the Pauli Exclusion Principle. This means that two indistinguishable fermions can never occupy the same quantum state.

Instead, in the limit of absolute zero temperature they fill the lowest energy states with exactly one atom per state in an arrangement known as the Fermi sea.

All energy states up to the Fermi energy E_F are filled with particles in two different spin states. Such a Fermi sea is not superfluid and the assembly of fermions will be effectively a non-interacting system of particles. Fermi gas forms a useful first approximation in the theory of metals, in the theory of ^3He , in studies of nuclear structure, nuclear matter and neutron stars.

To understand the behavior of such many-body assemblies composed of fermions, some interaction must be included between the particles. In this manuscript it is assumed that the interaction between two fermions in the system is a hard-sphere interaction.

It should also be understood that in the long wavelength limit only s-wave scattering is considered, and for hard-sphere gas the scattering length 'a' is just the hard-sphere diameter (Beliaev, 1958).

A hard-sphere system of fermions with density ranging from very low to very high will be considered to obtain an expression for the energy per particle $\left(\frac{E}{N}\right)$. For fermions, we need to satisfy Pauli exclusion principle which leads to repulsion between fermions when they try to approach each other to occupy the same energy state.

A gas or liquid composed of ^3He is a well known assembly of fermions in which particles can be assumed to interact via the pairwise potential (Aziz *et al.*, 1979), and the energy of such a system can be calculated. Neutron matter and nuclear matter are other systems of fermions.

For an N-identical fermion system, the energy E is given by (Baker, 1971) for a low-density system,

Research Article

$$\frac{E}{N} = \frac{3}{5} \frac{\hbar \kappa_F^2}{2m} \left\{ \begin{aligned} &1 + C_1(\kappa_F a) + C_2(\kappa_F a)^2 + \left[\frac{C_3 r_0}{2a} + \frac{C_4 A_1(0)}{a^3} + C_5 \right] (\kappa_F a)^3 \\ &+ C_6(\kappa_F a)^4 \log(\kappa_F a) + \left[\frac{C_7 r_0}{2a} + C_8 \frac{A_0''}{a^3} + C_9 \right] (\kappa_F a)^4 + 0(\kappa_F a)^4 \end{aligned} \right\} \quad 1$$

where the (Baker *et al.*, 1982) C_j 's ($j = 1, 2, \dots, 9$), for $\nu = 2$ and $\nu = 4$, are dimensionless co-efficients depending on ν , which is the number of intrinsic degrees of freedom of each fermion. For instance, for $\nu = 4$, $C_2 = 0.556610$ and C_9 is not available for $\nu = 4$. The Fermi-momentum $\hbar \kappa_F$ in terms of the fermion number density ρ is given by,

$$\rho = \frac{N}{V} = \frac{\nu \kappa_F^3}{6\pi} \quad 2$$

where, V is the volume of the system. The quantities a , r_0 and $A_1(0)$ are parameters containing information related to two-body scattering due to a central potential $U(r)$. Whereas, $A_0''(0)$ cannot be related to the scattering phase shift alone, but is potential-shape dependent and can thus be interpreted as the first correction to the static limit, while C_9 has a three-body cluster contribution.

In fact, the low density expansion given in Eqn (1) breaks down at moderate and high densities including the saturation liquid density of ^3He and nuclear matter.

For fermion hard-sphere system Eqn (1) can be re-written as,

$$\frac{E}{N} = \frac{3}{5} \frac{\hbar^2 \kappa_F^2}{2m} e_0(x) \quad 3$$

Where $x = \kappa_F \chi_0$ and χ_0 is the hard-core diameter. Now $e_0(x)$ can be written as,

$$e_0(x) \cong 1 + C_1 x + C_2 x^2 + \left(\frac{C_3}{3} + \frac{C_4}{3} + C_5 \right) x^3 + C_6 x^4 \log x + \left(\frac{C_7}{3} - \frac{C_8}{3} + C_9 \right) x^4 + 0(x^4) \quad 4$$

For $\nu = 2$, the value of (Baker, 1971; Baker *et al.*, 1982) $C_6 = 0$ and Eqn (4) simplifies to,

$$e_0(x) \cong 1 + D_1 x + D_2 x^2 + D_3 x^3 + D_4 x^4 + 0(x^4) \quad 5$$

for $x \ll 1$, with the D 's expressible in terms of the C 's. The ultimate aim is to calculate the value of E/N for very high densities when the system of hard-spheres of fermions will go to close packing (*cp*); it does not matter whether the packing is random or regular. But it should be understood that the regular *cp* will be face-centered-cubic, *fcc*, or hexagonal-close-packing, *hcp*.

Theory

Ultimately, the value of E/N has to be calculated from Eqn (3). This requires that the function $e_0(x)$ should not have a zero in its denominator since that will make the energy diverge. We have to work in the region of physical interest, and that is the one in which $0 \leq x \leq 3.47$, and in this region the energy does not diverge at any close packing. However, the value of $e_0(x)$ will be approximated as (Solis *et al.*, 2008),

$$e_0(x) \cong 1 + C_1 x + C_2 x^2 + \left(\frac{C_3}{3} + \frac{C_4}{3} + C_5 \right) x^3 \quad 6$$

For [4] $\nu = 2$ (two degrees of freedom for the fermions), $C_1 = 0.353$; $C_2 = 0.185$; $C_3 = 0.384$

Research Article

Here $x = \kappa_F \chi_0$ where $\chi_0 = 2.1117$ is the hard-core diameter for ^3He . Similarly, [5] for $v = 4$ (four degrees of freedom) $C_1 = 1.0610$; $C_2 = 0.5566$; $C_3 = 1.3006$.

The general expression for κ_F via the Fermi energy ϵ_F (Khanna, 1986) is given by,

$$\epsilon_F = \frac{\kappa_F^2 \hbar^2}{2m} = \left(\frac{\hbar^2}{8m} \right) \left(\frac{3}{\pi} \rho \right)^{\frac{2}{3}} \quad 7$$

or

$$\kappa_F = \left(3\pi^2 \rho \right)^{\frac{1}{3}} \quad 8$$

Where in our case the mass m is for ^3He atom, and $\rho = 0.86\rho_0$. But for cp or hcp , ρ in Eqn (8) may be replaced by $\rho_0 = \sqrt{2}/\chi_0^3$ where the assembly may have reached the crystalline state.

For hpc or fcc , the value of E/N is given by (Rhen and Llano, 1989),

$$\frac{E}{N} = Q_v \rho^{\frac{2}{3}} + \left(\frac{v-1}{v} \right) \frac{2\pi\hbar^2 \chi_0}{m} \left[\frac{1}{\left(\rho^{\frac{1}{3}} - \rho_0^{\frac{1}{3}} \right) \left(\rho^{\frac{1}{3}} + b(v)\rho_0^{\frac{1}{3}} \right)} \right] \quad 10$$

Where

$$Q_v = \frac{3\hbar^2}{10m} \left(\frac{6\pi^2}{v} \right)^{\frac{2}{3}} \quad 11$$

and

$$b(v) = \left[\left(\frac{v-1}{v} \right) (b+1) - 1 \right] \approx b \quad \text{for } v \rightarrow \infty \quad 12$$

With $v = N \rightarrow \infty$, in the low density limit,

$$\frac{E}{N} \underset{\rho \rightarrow 0}{=} Q_v \rho^{\frac{2}{3}} + \left(\frac{v-1}{v} \right) \frac{2\pi\hbar^2 \rho \chi_0}{m} \quad 13$$

Another expression for energy per particle that has been used for ^3He is given in Eqn (3) (Panoff, 1990; Panoff and Carlson, 1989).

The value of the total energy E can be obtained from Eqn (13) since all the parameters in Eqn (13) are known. In the limit of small ρ ($\rho \rightarrow 0$), let E be represented by E_0 . To understand the temperature dependence of the energy E , specific heat C_v , entropy S and the transition temperature T_c ; E_0 will be multiplied by the thermal activation factor τ , such that,

$$\tau = \exp \left(- \frac{E_0}{\kappa_B T} \right) \quad 14$$

where κ_B is the Boltzmann constant.

And the total energy E may be written as,

Research Article

$$E = E_0 \exp\left(-\frac{E_0}{\kappa_B T}\right) \quad 15$$

The specific heat C_v becomes,

$$C_v = \frac{\partial E}{\partial T} = \frac{E_0^2}{\kappa_B T^2} \exp\left(-\frac{E_0}{\kappa_B T}\right) \quad 16$$

The entropy S becomes,

$$S = \int \frac{C_v dT}{T} = \int \frac{E_0^2}{\kappa_B T^3} \exp\left(-\frac{E_0}{\kappa_B T}\right) dT \quad 17$$

or

$$S = \kappa_B \left(\frac{E_0}{\kappa_B T} + 1 \right) \exp\left(-\frac{E_0}{\kappa_B T}\right) \quad 18$$

The transition temperature T_c will be given by,

$$\left(\frac{\partial C_v}{\partial T} \right)_{T=T_c} = 0 = \frac{E_0^2}{\kappa_B T^2} \left(-\frac{E_0}{\kappa_B} \right) \left(-\frac{1}{T^2} \right) \exp\left(-\frac{E_0}{\kappa_B T}\right) + \left(\frac{2E_0^2}{\kappa_B T^3} \right) \exp\left(-\frac{E_0}{\kappa_B T}\right) \quad 19$$

For $T = T_c$ in Eqn (19), we get,

$$T_c = \frac{E_0}{2\kappa_B} = 18.67 K \quad 21$$

T_c will be the temperature at which phase transition can take place in a system of hard-sphere gas of fermions (^3He).

Calculations

To do calculations for a system composed of ^3He atoms, $v=2$, Q_v is given by Eqn (11), $\rho = 0.86 \frac{\sqrt{2}}{\chi_0^3}$ but

$\chi_0 = 2.1117 \text{\AA}$, and assuming N to be the number of particles in the unit volume, then $N = \rho$.

The mass of ^3He atom $m = 3.0160293 \times 10^{-24} \text{g}$,

Eqn (13) will be used to get the value of E_0 and Eqn (15) for E . Specific heat C_v will be calculated from Eqn (16), entropy S from Eqn (18) and the transition temperature T_c using Eqn (21).

Research Article

RESULTS AND DISCUSSION

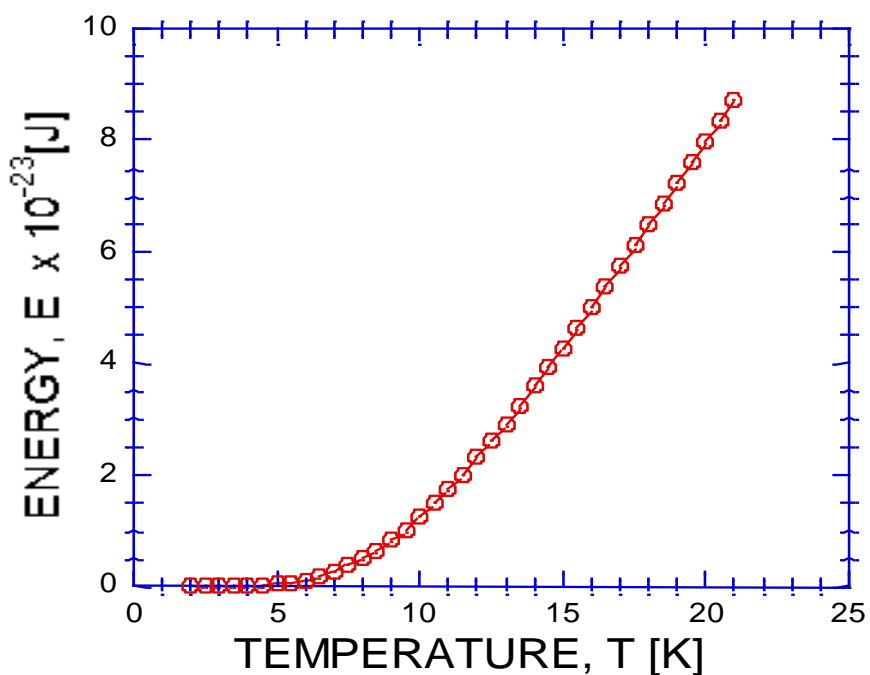


Figure 1: Energy variation with Temperature

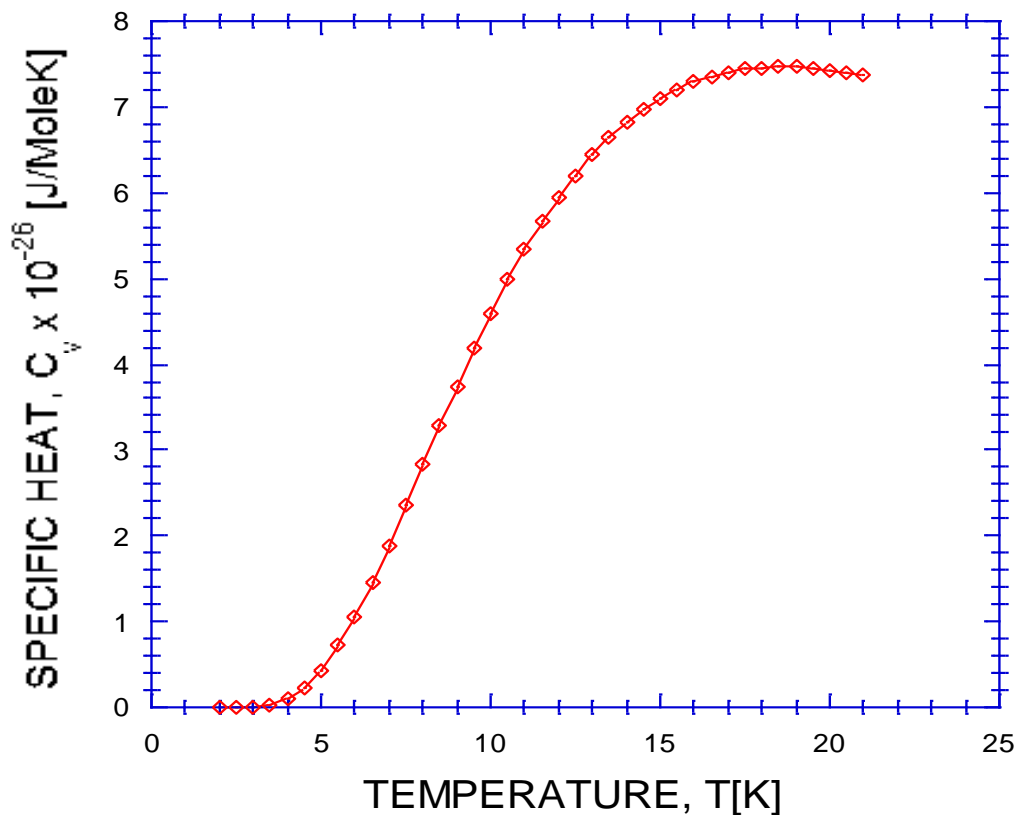


Figure 2: Specific Heat variation with Temperature

Research Article

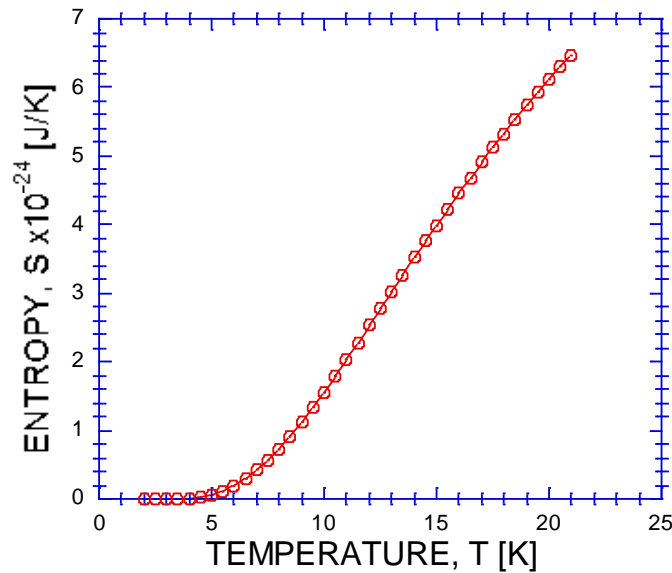


Figure 3: Entropy variation with Temperature

The calculations done for E , C_v and S are plotted in Figure 1, Figure 2 and Figure 3 respectively. Figure 1 shows that the energy E of the system increases as T increases, and this is what it should be thermodynamically. Figure 2 shows the variation of C_v with T , and as T increases, the value of C_v shows a point of inflection at a temperature T_c called the critical or transition temperature, and the value of $T_c = 18.67$ K. According to this calculation, a gas of ^3He will undergo a phase transition at $T_c = 18.67$ K provided the ^3He atoms are assumed to be interacting via a hard-sphere interaction. Such an interaction is characterized by the s-wave scattering length “a” (hard-core diameter $= \chi_0$). Mean-field theory gives a relation (Holland *et al.*, 2001; Leggett, 1980; Heiselberg, 2000) between T_c and the Fermi temperature T_F such that,

$$\frac{T_c}{T_F} \approx \exp\left(-\frac{\pi}{2|\chi_0|\kappa_F}\right) \quad 19$$

The Fermi temperature for fermions is,

$$T_F = \frac{\hbar^2}{2m\kappa} (3\pi^2 \rho)^{\frac{2}{3}} \quad 20$$

Where the momentum vector, $\kappa_F = (3\pi^2 \rho)^{\frac{1}{3}}$, and giving the value $|\chi_0|\kappa_F = 3.30$. Since for the ^3He hard-sphere gas, the values of $|\chi_0|\kappa_F$ may lie between 0 and 3.47, in Eqn (19) using $|\chi_0|\kappa_F = 3.30$, the value of $T_c = 20.3$ K. This transition temperature closely compares with our calculated value of 18.67K. The variation of C_v with T in Fig. 2 qualitatively agrees with the variation of C_v with T (Mishra and Ramakrishnan, 1985).

Our calculations lead to the conclusion that an assembly of hard-sphere ^3He can undergo a phase transition at a transition temperature $T_c = 18.67$ K.

REFERENCES

- Aziz RA, Nain VPS, Carley JS and Taylor WL (1979).** An accurate intermolecular potential for helium. *McConville, G.T., Journal of Chemical Physics* **70** 4330.
Baker Jr GA, Benofy LP, Fortes M, Peltier S and Plastino A (1982). Hard-core square-well fermion. *Physical Review* **A26** 3575.

Research Article

Baker Jr. GA (1971). Singularity Structure of the Perturbation Series for the Ground-State Energy of a Many-Fermion System. *Review of Modern Physics* **43** 479.

Beliaev ST (1958). Application of the methods of quantum field theory to a system of bosons. *Soviet Physics JETP* **7** 299.

Heiselberg H, Pethik CJ, Smith H and Viverit L (2000). Influence of Induced Interactions on the Superfluid Transition in Dilute Fermi Gases. *Physical Review Letters* **85** 2418.

Holland M, Kokkelmans SJJMF, Chiofalo ML and Walser R (2001). Resonance Superfluidity in a Quantum Degenerate Fermi Gas. *Physical Review Letters* **87** 120406.

Khanna KM (1986). *Statistical Mechanics and Many-Body Problem* (To-day and To-morrow Printers and Publishers) New Delhi.

Leggett AJ (1980). Modern Trends in the Theory of Condensed Matter. *Journal of Physics (Paris)* **41**(C7) 19.

Mishra SG and Ramakrishnan TV (1985). Fluctuation Theory of Specific Heat of Normal Liquid ^3He . *Physical Review* **B31** 2825.

Panoff RM (1990). *Recent Many-Body Theories* (Plenum) New York **2**.

Panoff RM and Carlson J (1989). Fermion Monte Carlo algorithms and liquid ^3He . *Physical Review Letters* **62** 1130.

Rhen SZ and Llano MDE (1989). Generalised London equation of state for fermion hard-sphere fluids. *European Journal of Physics* **10**(2) 96.

Solis MA, Llano MDE, Clark JW and Baker Jr GA (2008). Improved Quantum Hard-Sphere Ground-State Equations of State. *Condensed-matter-statistical-mechanics*.